High Spectral/Temporal Resolution Cloud Detection

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Identify cloud and surface characteristics in high spectral resolution data which best delineate clouds, aerosols, and surface properties from one another

Develop a cloud detection algorithm that exploits high spectral resolution measurements such as GIFTS/IOMI

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Past Year Accomplishments

UAH student (Kevin Laws) complete M.S. refining cloud detection algorithm with GOES – BTH technique – <u>demonstrates value of spatial</u> <u>and temporal thresholds in difference images</u>



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Past Year Accomplishments (continued)

New UAH graduate student (Nicole Slodysko) began thesis work with analysis of AIRS data

- analyzed AIRS spectrum of clouds
- categorized spectral signatures day, night, cloud type, view angles
- preliminary insight on detection strategy



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Earth-Atmosphere Emission Spectrum

Examine Earth-atmosphere emission spectrum (AIRS)

• Relatively transparent (window) regions detect emission/reflective properties from surface and clouds

• GOES and MODIS look at this emission in relatively broad bands – measure integrated effect of spectral emission



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Spectral Signature of Clouds

<u>GOES</u> BTH technique: 11μ m, and $11-3.9 \mu$ m differences (day and night)

• spatial and temporal thresholds determined

EOS approach with MODIS uses many spectral difference tests

11–3.9 $\mu m, ~~11$ -12 $\mu m, ~~8.6$ -11 $\mu m, ~~3.7$ -12 $\mu m, ~~11$ -6.7 μm

These approaches can be successful but do not fully utilize high spectral information!



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High Spectral Resolution Methodology

Examine high resolution spectral differences in AIRS data for cloud signatures

Stratify AIRS scenes into categories

- Day and night
- Cloudy and clear
- Types of clouds (cirrus of vary optical depth, various levels of opaque clouds)

Calculate average of 10 AIRS channels centered on 11.08 μ m (T_{bb}(11)) (between H₂O absorption lines)

Calculate and examine difference spectrum

Use results in spectral cloud detection scheme

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T_{bb} (11) – T_{bb} (3-4)µm Difference Spectrum for Clouds

Brightness Temperature Over Clear Alabama Skies

Brightness Temp. BT11 Diff.

DTTO:

NU.

-15

clear

300

280

260

¥

the state

Cloud (day):

- difference is (large) negative magnitude depends reflected solar
- hump at 3.9μm is missing
- 3.9-4.1µm slope is more negative
- 3.7-3.9µm slope significant

Cloud (night)

- difference is (small) positive (absence of reflected solar)
- 3.9-4.1µm slope is flat



T_{bb} (11) – T_{bb} (10-13)µm Difference Spectrum for Clouds

Cloud (day & night):

- negative slope more pronounced for cirrus clouds
- magnitude of absorption lines are small (little moisture above cloud)

cloudy

Brightness Temperature Over Cloudy Indiana Skies

Brightness Temp.

GT1.1 DWF

16 April 2003 19,30 UTC

10.5

- spectral variability is reduced
- no day-night differences

320

300

¥ andore

一日日日

260

240

May 28, 2003

13.0

12.5



115

Wavelength (micron)

11.0

12.0

T_{bb} (11) – T_{bb} 11(8-9)µm Difference Spectrum for Clouds

Cloud (day & night):

- negative differences for cirrus clouds
- magnitude of absorption lines are small (little moisture above cloud)

cloudy

8.2

- spectral variability is reduced
- no day-night differences

3000

290 F

 Σ

270

260

250

240

3.4



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Preliminary Findings

High spectral resolution measurements in the short and long wave infrared regions provide significant information for the detection of clouds beyond rather broad channel measurements (like those of MODIS and GOES)

• Eliminate effects of absorption lines – changing amplitude of lines useful as well

• Spectral difference plots show slope and offset differences useful to detect presence and type of clouds



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Future Work

Apply slope and offset methodology to develop a cloud detection and identification scheme for high spectral resolution data

- use multiple spectral difference plots to develop spatial/temporal slope and intercept images
- apply to high-resolution data for cloud retrieval
- explore AIRS channel-tochannel cloud signatures
- evaluate 5.0 μm window region in GIFTS cubes and a/c data



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