



# UV Remote Sensing of Volcanic Ash

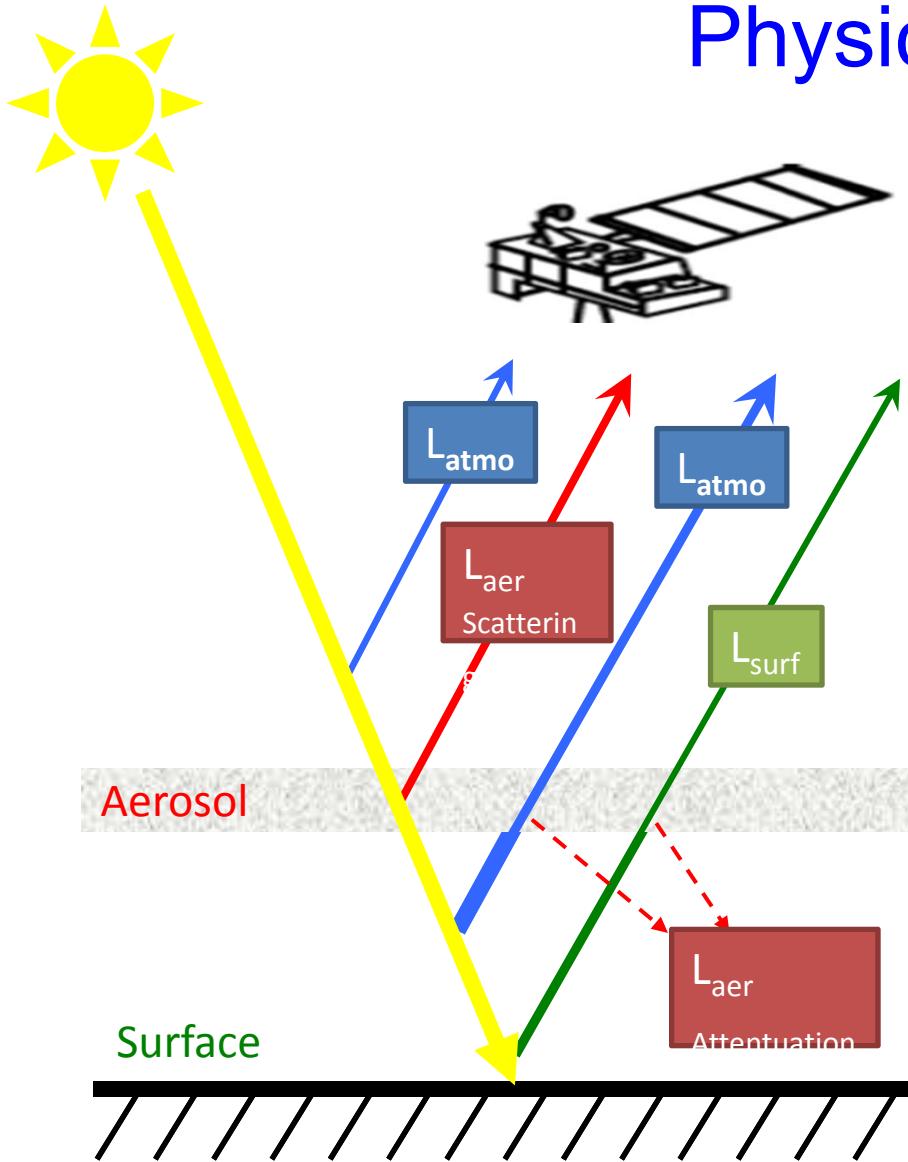
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WMO Inter-comparison of Satellite-based Volcanic Ash Retrieval Algorithms Workshop

June 26 – July 2, 2015, Madison, Wisconsin

# Remote Sensing of Aerosols: Physical Basis



The top-of-atmosphere radiance ( $L_{TOA}$ ) can be separated into three different contributions for an aerosol laden atmosphere.

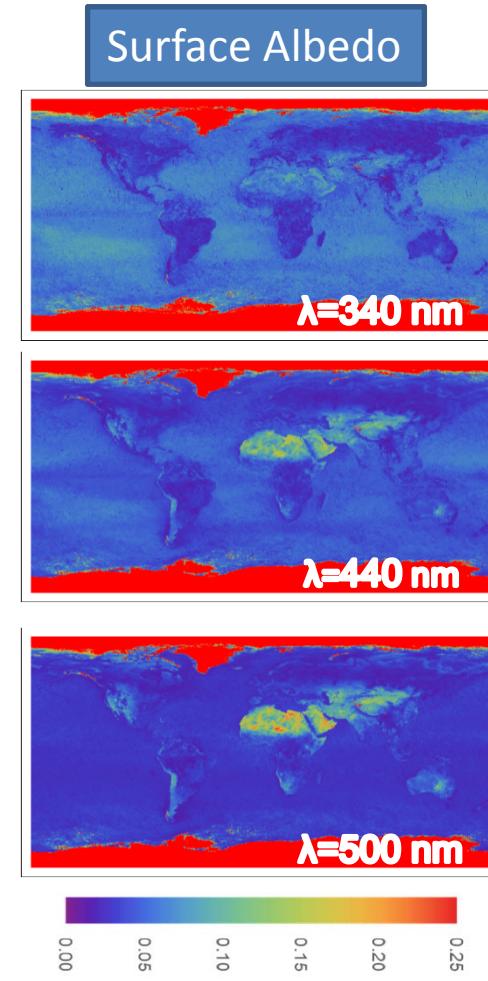
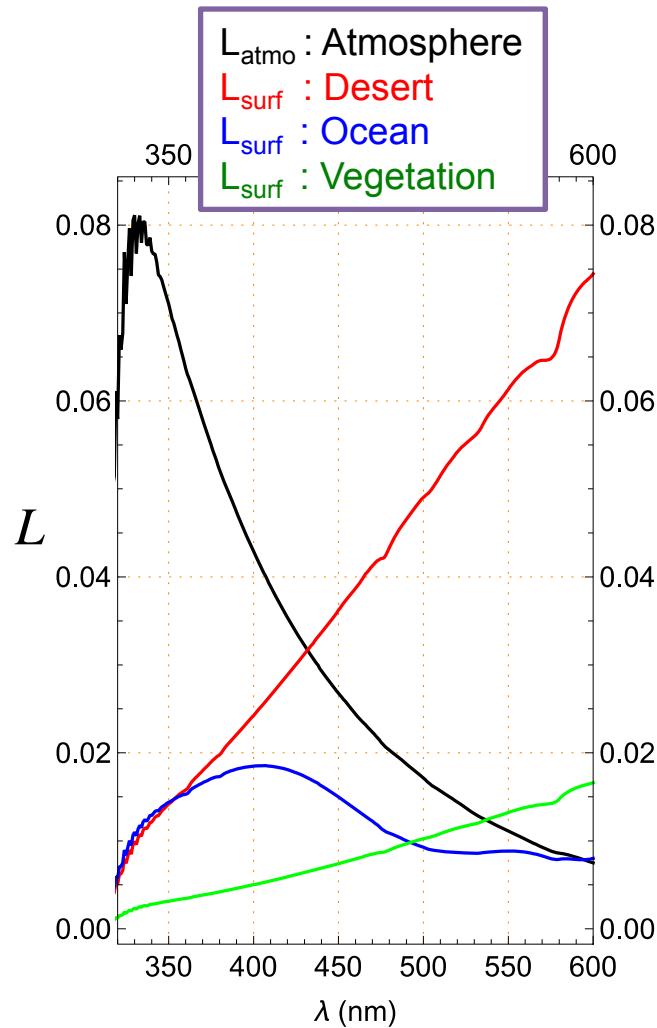
Backscattered radiance by air molecules:  
Rayleigh scattering +  
gas absorption

Radiance change due to aerosol scattering and absorption

$$L_{TOA} = L_{atmo} + L_{surf} + L_{aer}$$

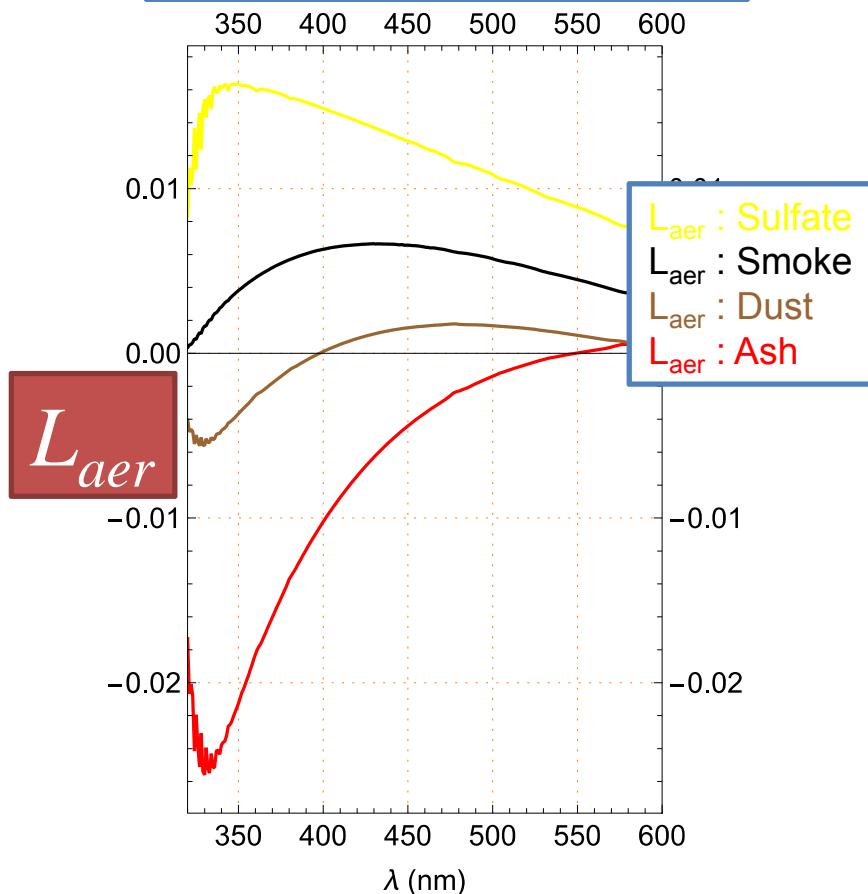
Radiance reflected from underlying surface

# Spectral Dependence of Radiance contributions: Atmospheric Backscattering and Surface Reflection



# Spectral Dependence of Aerosol Effects: $L_{aer}$

$L_{aer}$  : Radiance change due to aerosols



$$L_{aer}(\lambda) \approx AS + (L_{surf} + L'_{atmo})(1-AA)$$

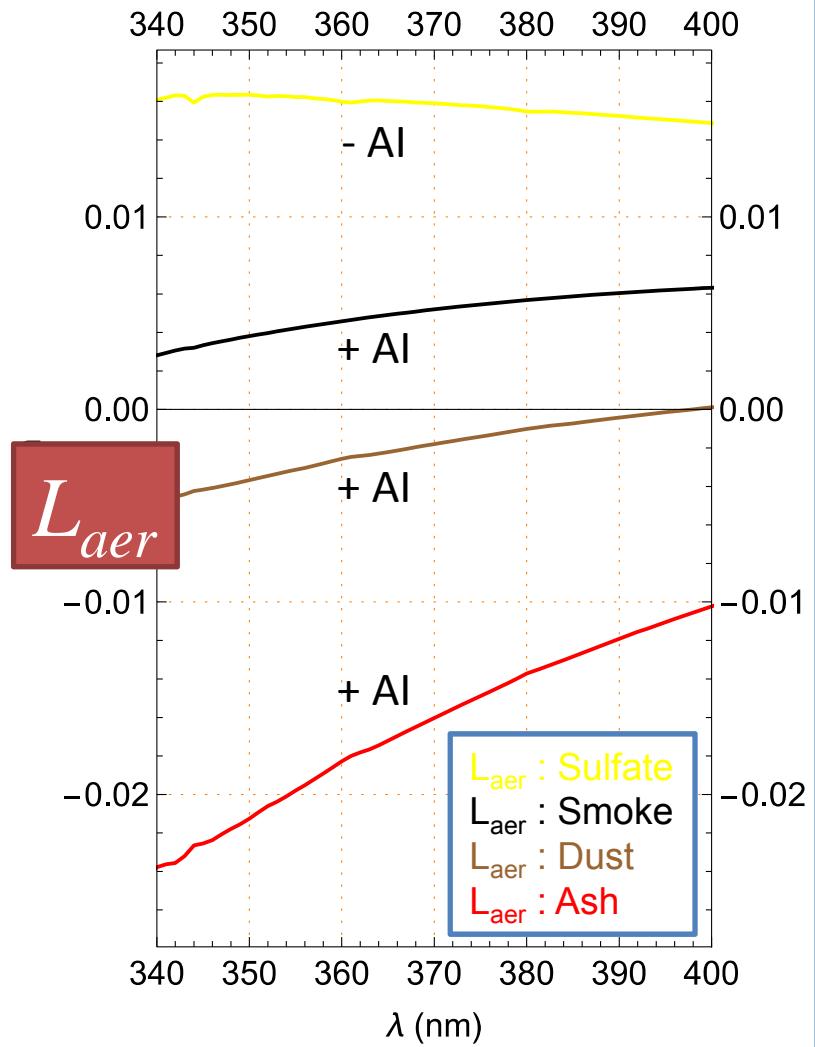
where AS is aerosol scattering,

AA is aerosol absorption,

$L'_{atmo}$  is atmospheric radiance from under the aerosol layer.

- $L'_{atmo}$  is large in UV, small in VIS/NIR.
- In UV, aerosol measurement is accomplished by quantifying its scattering and absorption effects.
- In VIS/NIR, aerosol measurement is primarily relied on quantification of aerosol scattering, when surface is dark or when surface reflection is properly accounted for.

# UV Aerosol Index (AI): Quantification of Radiance Change



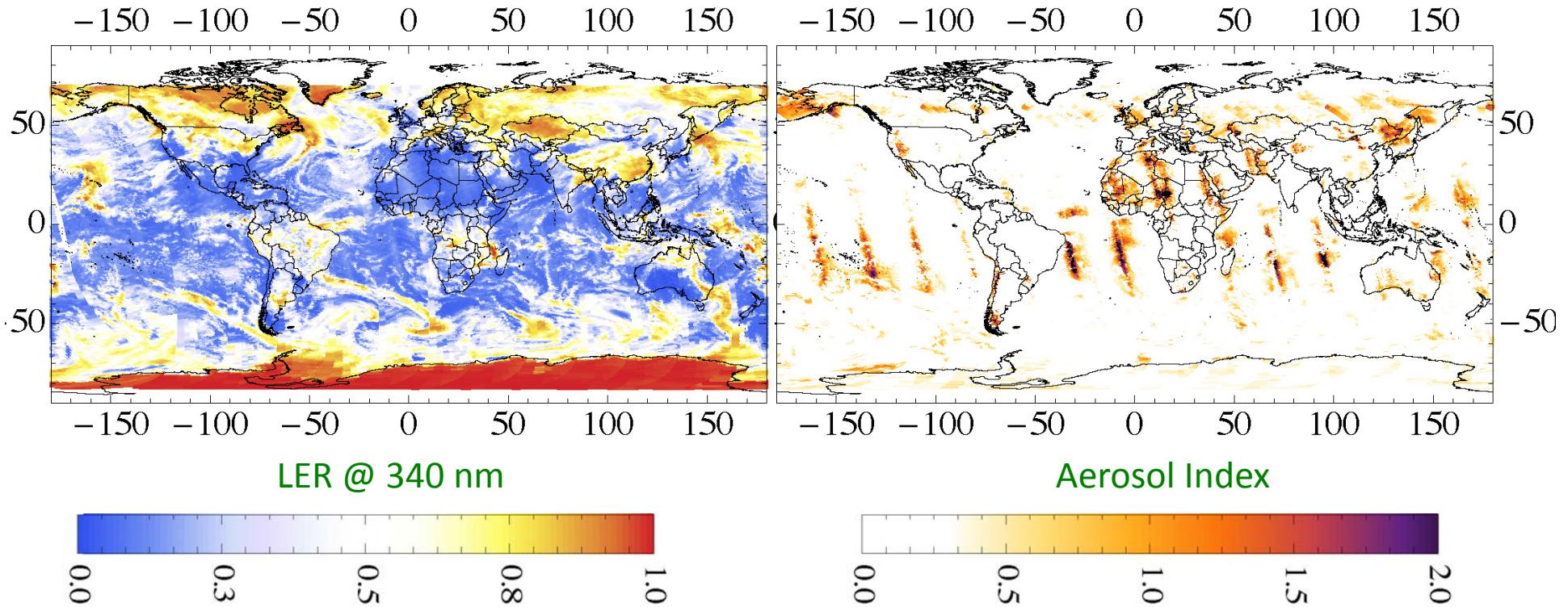
## Spectral Contrast in Apparent Surface Reflectivity $R_\lambda$

- Definition of Aerosol Index (AI): Spectral slope of  $R_\lambda$ , proportional to AI value.
- The spectral dependence of TOA radiance change is most pronounced for UV-absorbing aerosols, which cause  $R_\lambda$  to increase with wavelength: +AI
- Non-absorbing aerosols, under certain conditions, can cause  $R_\lambda$  to decrease with wavelength: - AI

# UV Aerosol Index (AI)

- AI is computed without any information about the aerosol particles (e.g., the refractive index and particle size distribution), and is determined by the deviation from Rayleigh atmosphere.
- AI can be used to determine their location and the relative amount of UV absorbing aerosols, even over bright surfaces, such as snow/ice or meteorological clouds.

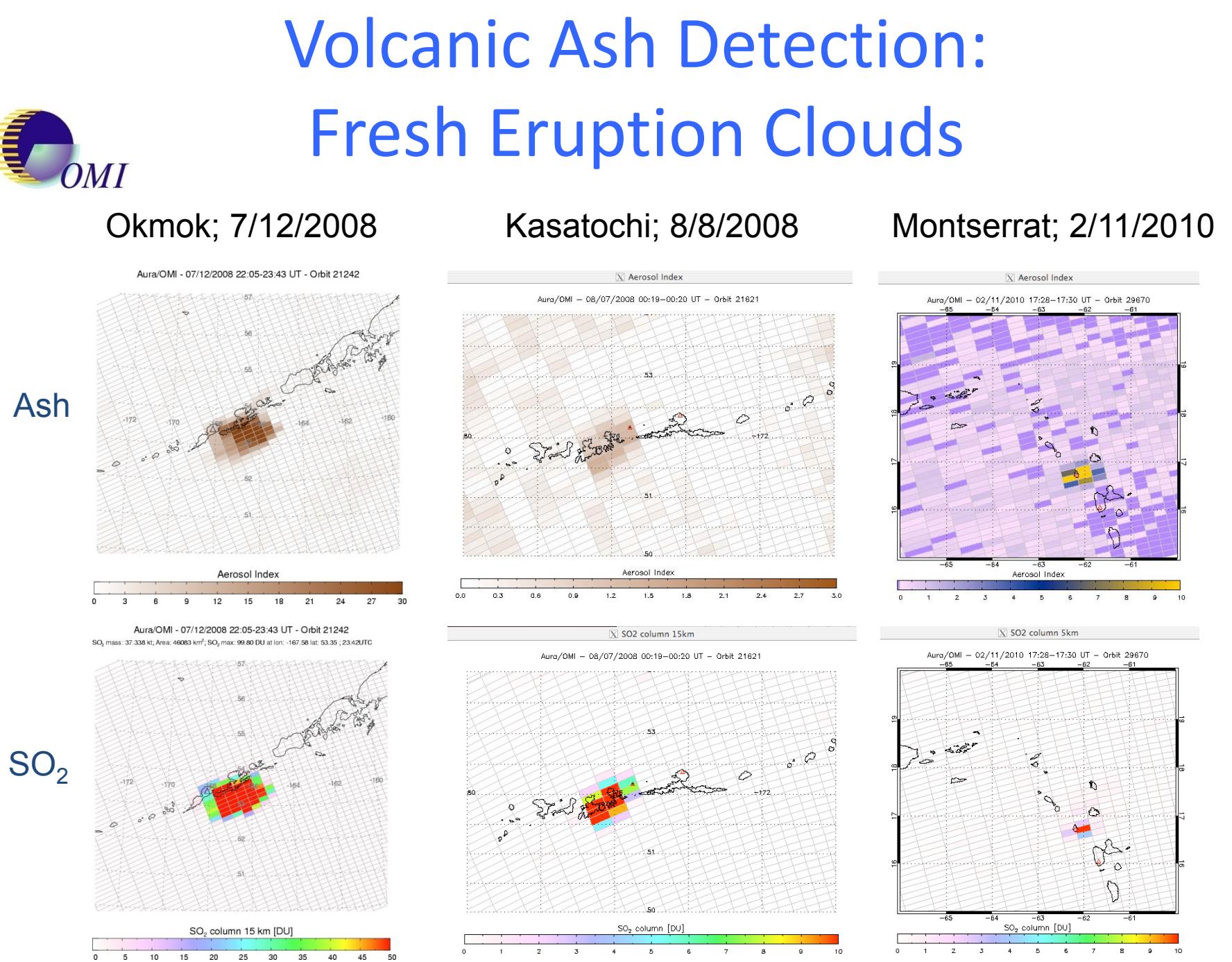
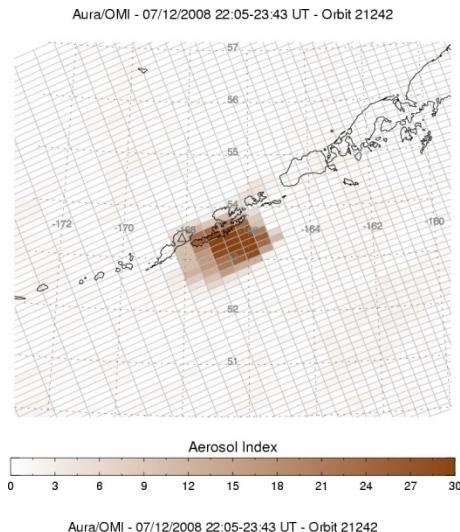
# Mapping of UV Absorbing Aerosols: Sample AI Data from SNPP/OMPS



# Volcanic Ash Detection: Fresh Eruption Clouds



Okmok; 7/12/2008

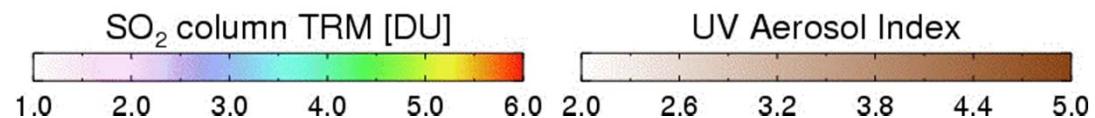
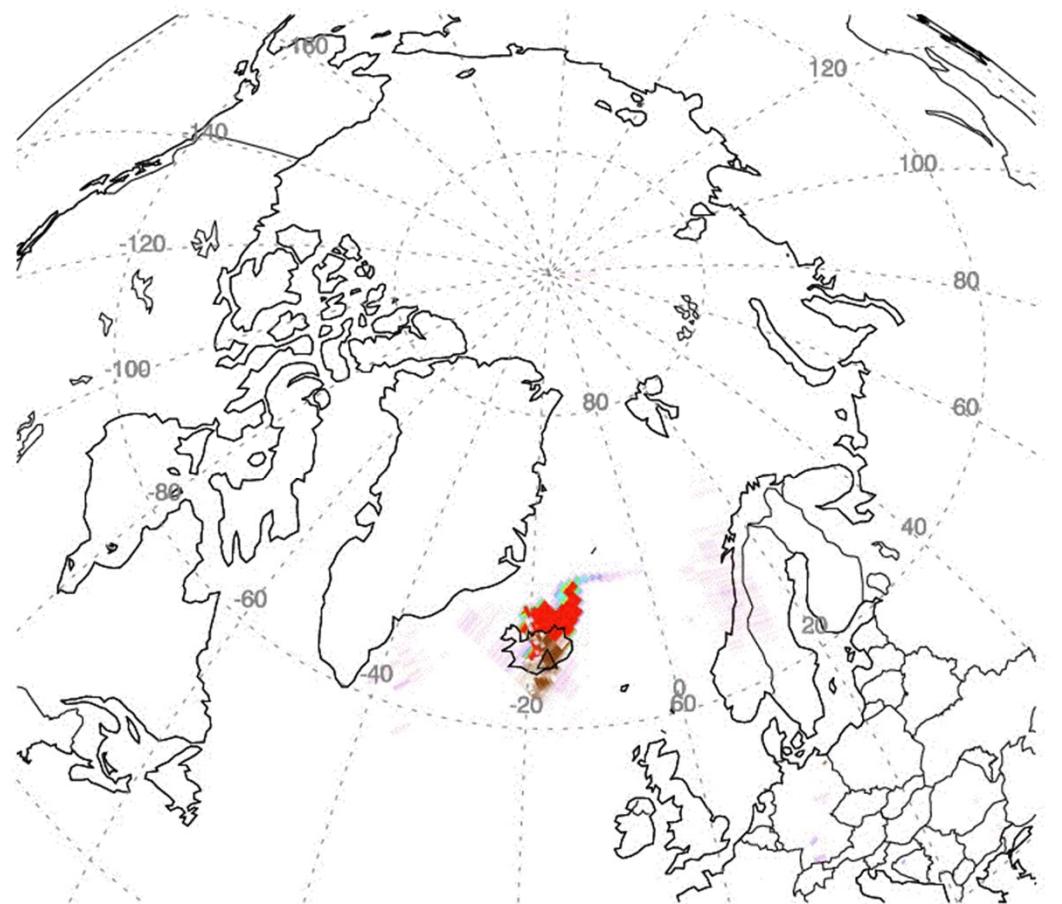


# Ash and SO<sub>2</sub>: Grímsvötn, May 2011



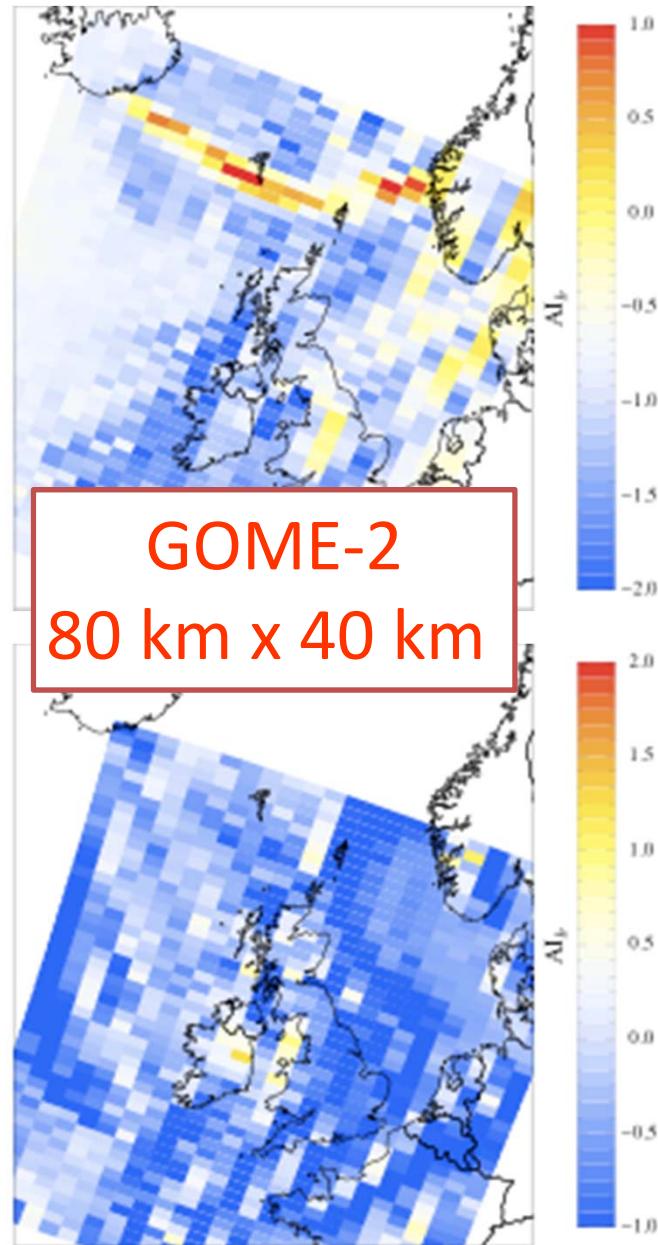
Aura/OMI - 05/22/2011 11:09-13:33 UT

SO<sub>2</sub> mass: 147.31 kt; Area: 197116 km<sup>2</sup>; SO<sub>2</sub> max: 241.38 DU at lon: -18.99 lat: 65.53 ; 13:27UTC

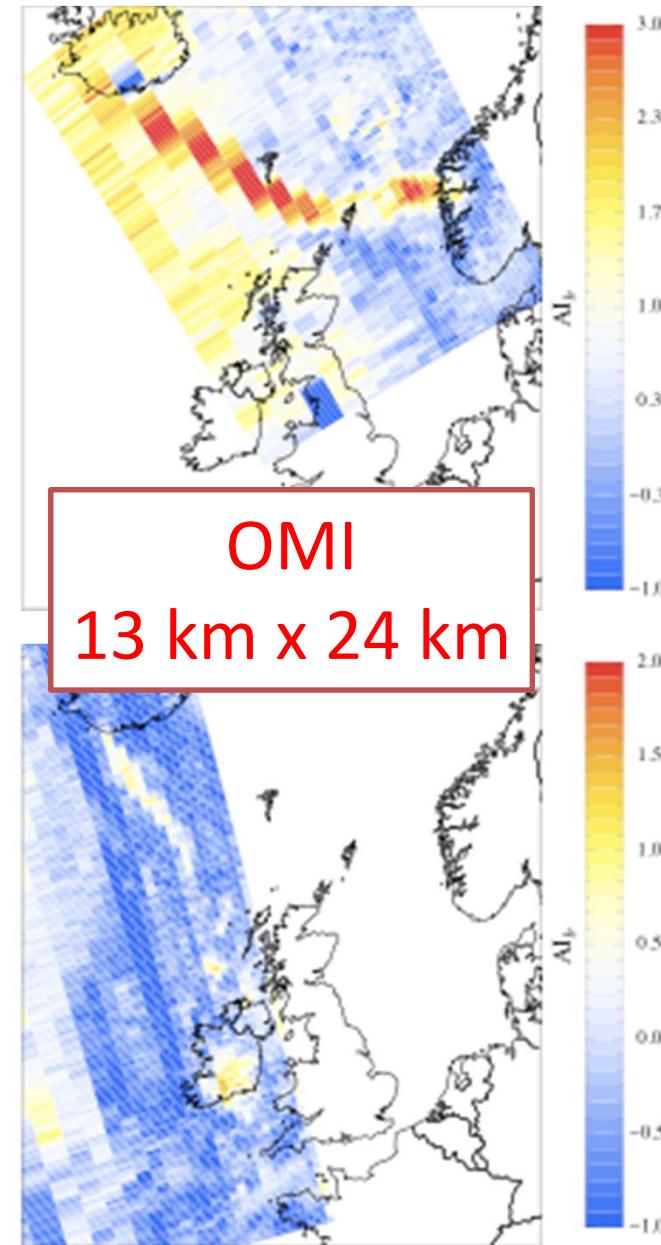


# Eyjafjallajökull Ash: Pixel Size Effect

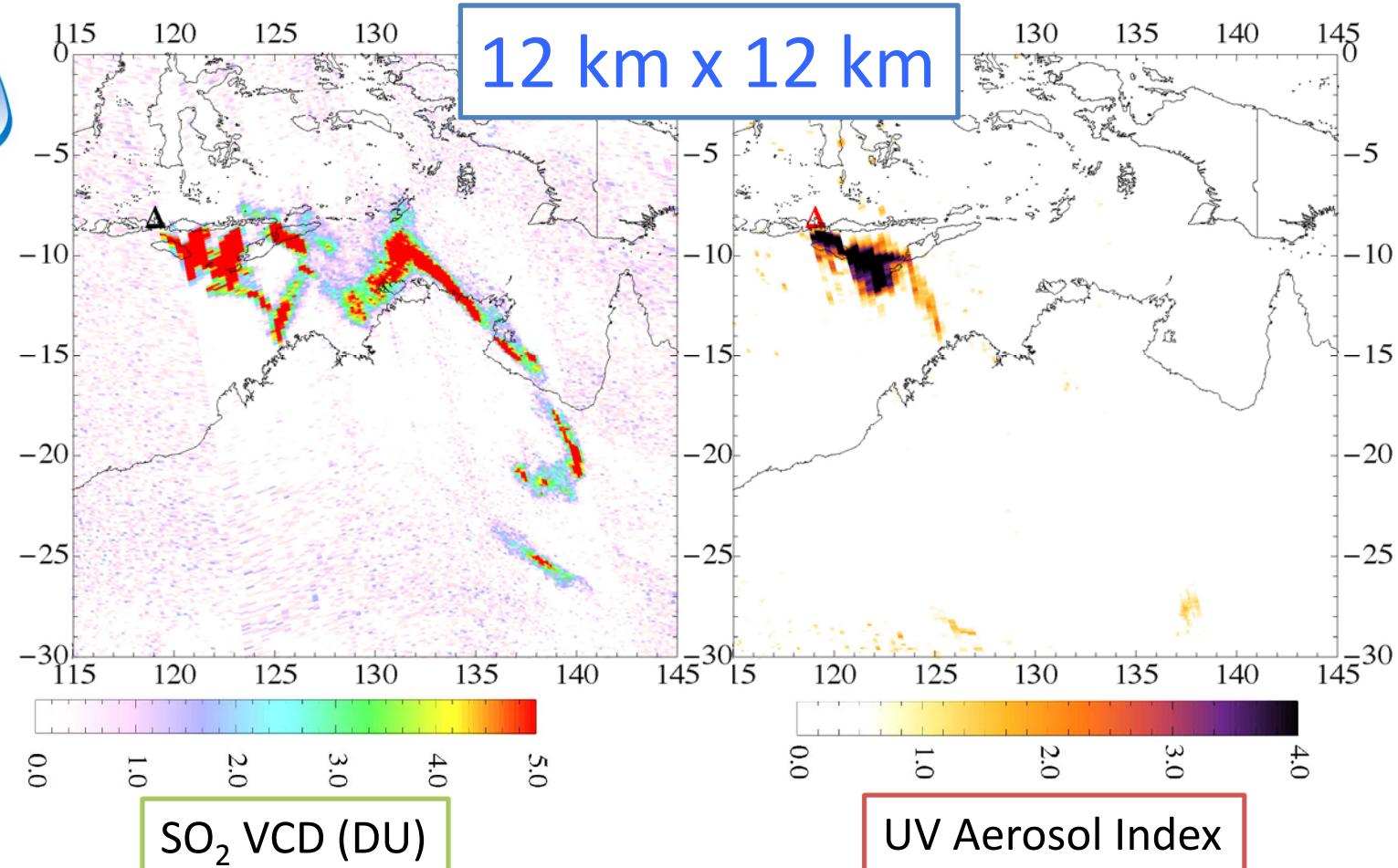
April 15, 2010



May 5, 2010



# SNPP/OMPS @ High Spatial Resolution Mode Future JPSS-1&2 will have similar high resolution

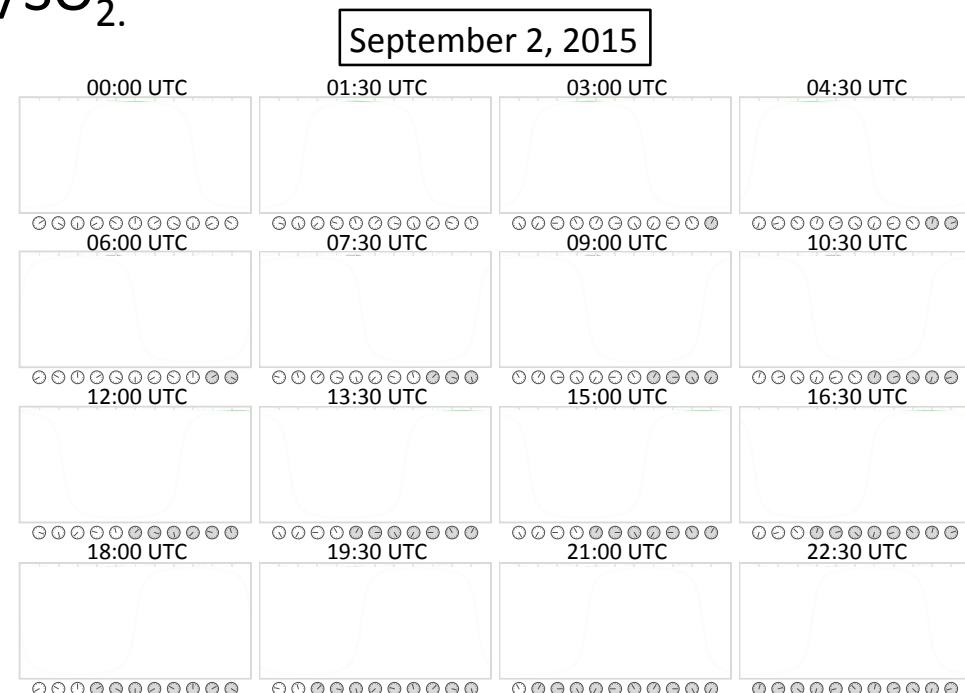


Eruption of Sangeang Api (Indonesia) 05/31/2014

# Near Future Enhancements from New Satellite Mission: DSCOVR (2015)



EPIC on NOAA/NASA DSCOVR at L-1 point, observes sunlit side of the Earth (UV/VIS discrete channels) at a spatial resolution of 24 km x 24 km, provides ash/SO<sub>2</sub>.

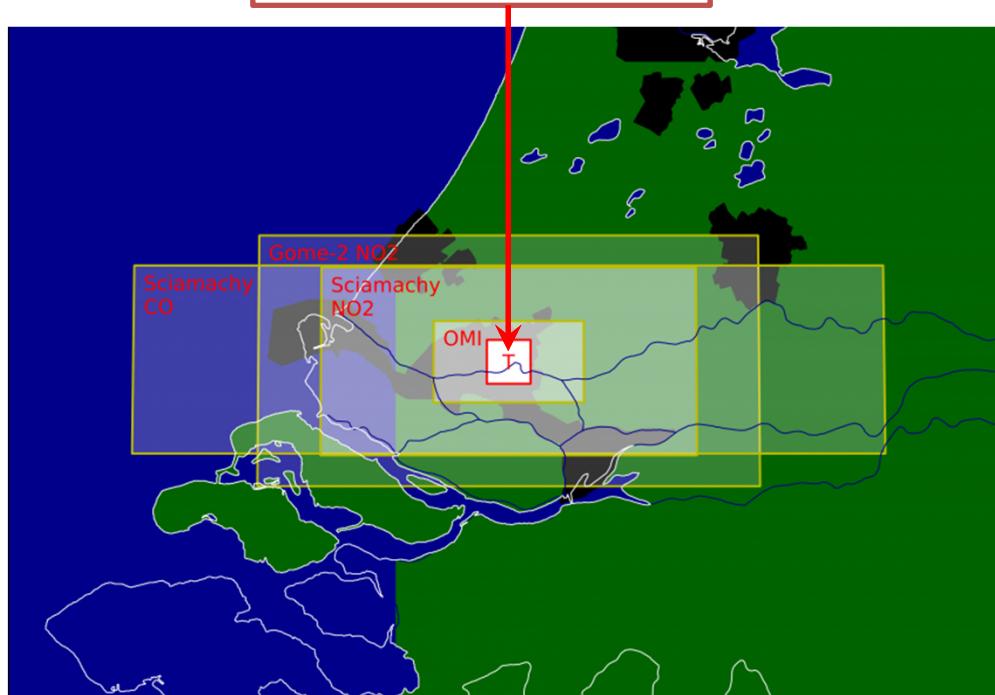


Multiple (>= 5) views in 24 hours

# Near Future Enhancements from New Hyperspectral UV/VIS Satellite Mission: TROPOMI (2016)



- TROPOMI on ESA polar orbiting Sentinel-5 Precursor, provides ash/SO<sub>2</sub> at a spatial resolution of **7 km x 7 km**.



# Quantification of Volcanic Ash

## Volcanic Cloud: Mixtures of Water/Ice Clouds and Ash Particles

Micro-physical properties of each component:

Particle shape (e.g. sphere/spheroid/irregular), mass density (e.g.,  $\rho_{\text{ash}} = 2.75 \text{ g/cm}^3$ )

Size distribution (e.g., Log-normal for ash particles,  $r_{\text{eff}} = 2\mu\text{m}$ ,  $\sigma = 1.6$ ),

Refractive index (e.g. real = 1.5, imag = 0.005, independent of  $\lambda$ )



## Volcanic Cloud: Optical Properties

Mass Coefficients: Scattering ( $K_{\text{sca}}$ ), Absorption ( $K_{\text{abs}}$ ),  
Extinction ( $K_{\text{ext}} = K_{\text{sca}} + K_{\text{abs}}$ ), Single Scattering Albedo ( $\omega_0 = K_{\text{sca}} / K_{\text{ext}}$ ),  
and Scattering Phase Function

## Vertical Distribution of Particles:

Layer height estimated from radiance measurements or trajectory analysis

## Surface Albedo

Estimated from radiance measurements or based on climatology



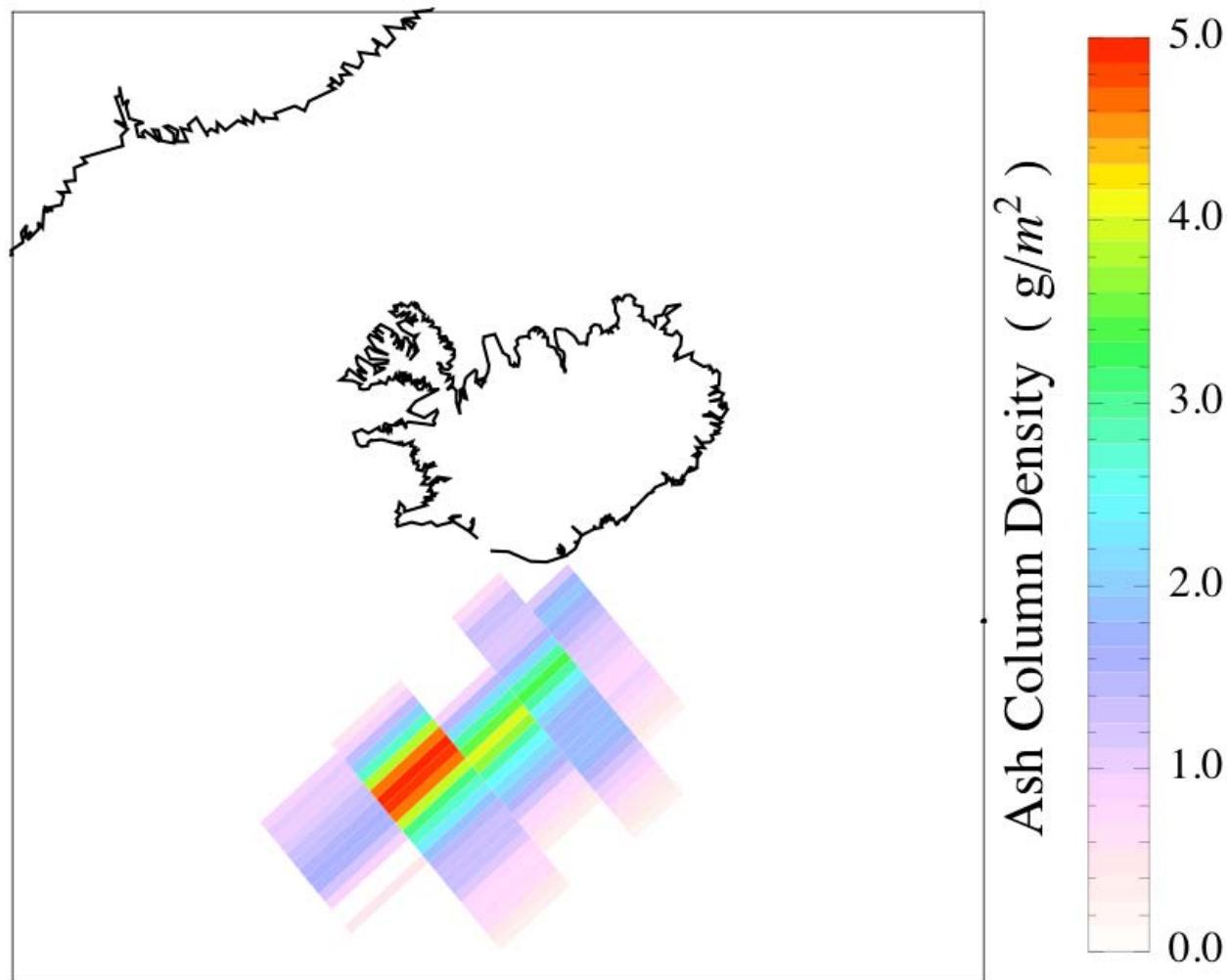
Extinction Optical Depths for Each Components: estimated from satellite radiance spectra



Ash Mass Concentration ( $\text{g/m}^2$ )

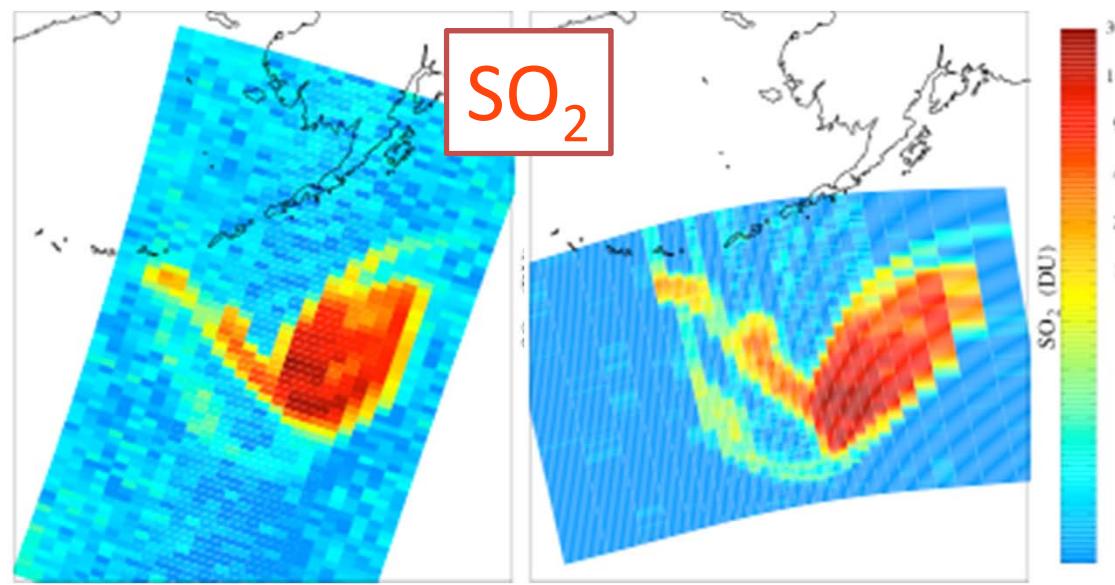
# Ash Concentration Estimation

Grimsvötn, 05/23/2011 OMI



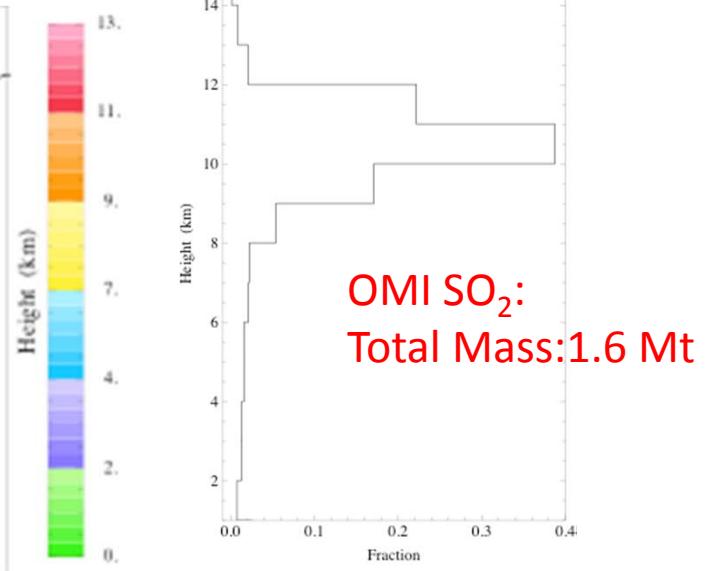
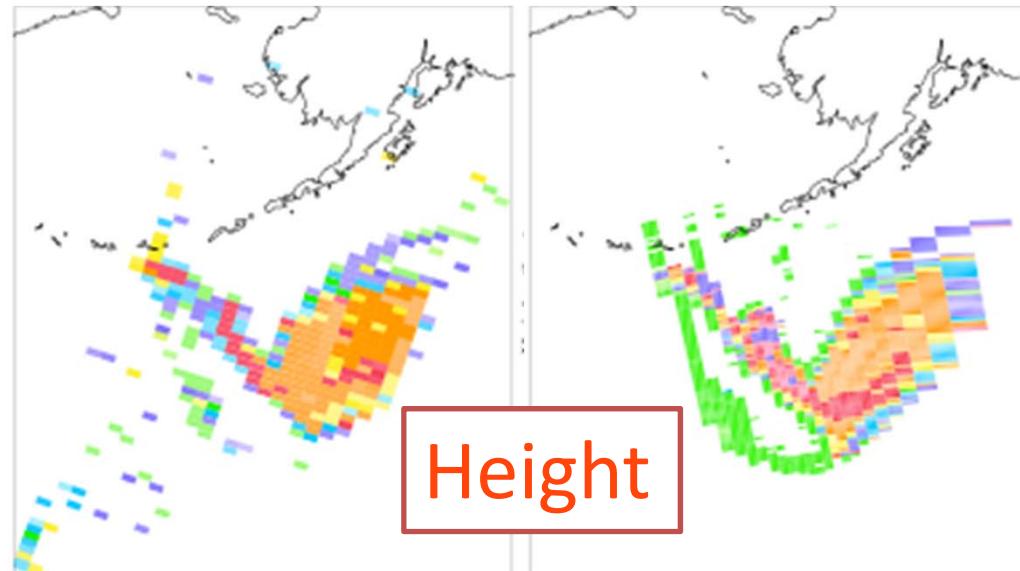
Kasatochi SO<sub>2</sub> Plume: August 9, 2008

# Plume Height Retrieval



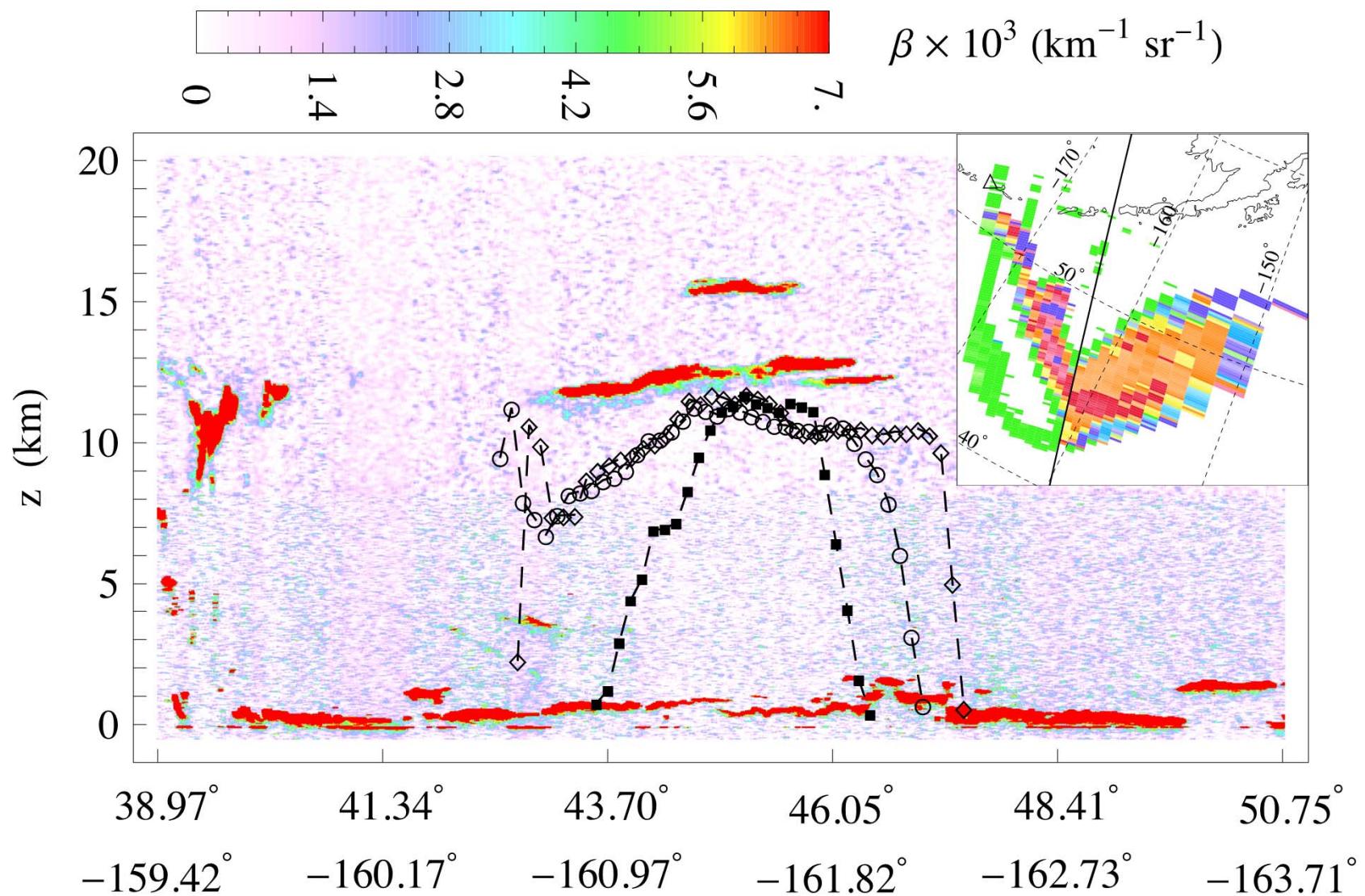
GOME-2

OMI



# August 2008 Kasatochi Eruption

## Comparisons with CALIPSO



# Value of UV Data

UV spectra are highly sensitive to ash (absorbing aerosols) and SO<sub>2</sub> in the atmosphere.

Volcanic clouds under a wide range of conditions:

- Detectable independent of water/ice content or surface conditions
- Detectable for fresh (dense) plumes
- Detectable for aged (weak) SO<sub>2</sub> plumes: long-term tracking
- Detectable down to the lower troposphere, including SO<sub>2</sub> from degassing: volcanic unrest
- Plume height from SO<sub>2</sub> measurements
- Ash amount given ash cloud particle properties

# Synergy of UV and IR

- Both UV and IR measurements are sensitive to ash particle size and composition, and its vertical location.
- Combining hyper-spectral UV (OMI, GOME2, OMPS) and IR (AIRS, IASI, CrIS) measurements provides greater constraints to a retrieval algorithm, and likely leads to more accurate estimates of volcanic ash particle size, plume height and loading.